

Math 481

17. $\int_{\gamma} \alpha$.

• We want to define $\int_{\gamma} \alpha$, for any r -form α and singular r -cube γ in a manifold M :

a) Definition: In coordinates $(U, (x^1, \dots, x^n))$ on M , we have $\gamma : [0, 1]^r \rightarrow M$, $\gamma(u^1, \dots, u^r) = (x^1(u^1, \dots, u^r), \dots, x^n(u^1, \dots, u^r))$, and $\alpha = \sum_{\vec{I}=(1 \leq i_1 < \dots < i_r \leq n)} a_{\vec{I}} dx^{\vec{I}}$. So

$$\begin{aligned} \gamma^* \alpha &= \sum_{\vec{I}} a_{\vec{I}} \left(\sum_{j_1=1}^r \frac{\partial x^{i_1}}{\partial u^{j_1}} du^{j_1} \right) \wedge \dots \wedge \left(\sum_{j_r=1}^r \frac{\partial x^{i_r}}{\partial u^{j_r}} du^{j_r} \right) \\ &= \sum_{\vec{I}} a_{\vec{I}} \det \frac{\partial (x^{i_1}, \dots, x^{i_r})}{\partial (u^1, \dots, u^r)} du^1 \wedge \dots \wedge du^r \end{aligned}$$

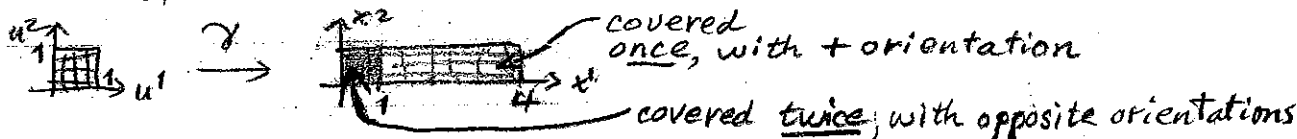
So, now that we've written $\gamma^* \alpha = f(u^1, \dots, u^r) du^1 \wedge \dots \wedge du^r$, just define

$$\int_{\gamma} \alpha = \int_0^1 \dots \int_0^1 f(u^1, \dots, u^r) du^1 \dots du^r.$$

b) Example: Line integrals, see Handout #15 d).

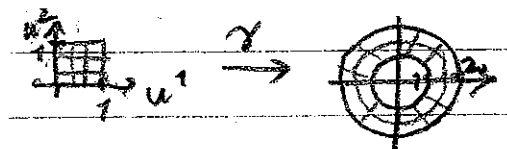
c) Example: $M = \mathbb{R}^2$. Take the 2-form $\alpha = dx^1 \wedge dx^2$ (the Euclidean "area form") and the singular 2-cube $\gamma(u^1, u^2) = (9(u^1 - \frac{1}{3})^2, u^2)$, $0 \leq u^i \leq 1$. Then

$$\int_{\gamma} \alpha = 3.$$



d) Example: $M = \mathbb{R}^2$, $\alpha = dx^1 \wedge dx^2$, $\gamma(u^1, u^2) = ((u^1 + 1) \cos(2\pi u^2), (u^1 + 1) \sin(2\pi u^2))$.

$\int_{\gamma} \alpha$ should be $\pi(2)^2 - \pi(1)^2 = 3\pi$. Check this!



e) In d), $\alpha = d(\frac{1}{2}(x^1 dx^2 - x^2 dx^1)) = d\beta$, where $\beta = \frac{1}{2}(x^1 dx^2 - x^2 dx^1)$. So,

$$\int_{\gamma} \alpha = \int_{\gamma} d\beta = 3\pi.$$

Now, calculate $\int_{\partial\gamma} \beta$. We have

$$\partial\gamma = \gamma_{2,0} + \gamma_{1,1} - \gamma_{2,1} - \gamma_{1,0} = \gamma_{1,1} - \gamma_{1,0},$$

where

$$\begin{aligned} \gamma_{2,0}(u^1) &= (u^1 + 1, 0), \quad \gamma_{1,1}(u^2) = (2 \cos(2\pi u^2), 2 \sin(2\pi u^2)), \\ \gamma_{2,1}(u^1) &= (u^1 + 1, 0), \quad \gamma_{1,0}(u^2) = (\cos(2\pi u^2), \sin(2\pi u^2)). \end{aligned}$$

Check that the answer is ... 3π !

4.6. Maxwell's Field Equations

In classical electromagnetic field theory one deals with the following quantities:

\mathbf{E} = electric field	\mathbf{H} = magnetic field
\mathbf{B} = magnetic induction	\mathbf{J} = electric current density
\mathbf{D} = dielectric displacement	ρ = charge density.

These are all functions of the space variables x^1, x^2, x^3 and the time t . The basic Maxwell equations in ordinary vector language are

- (i) $\text{curl } \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t}$ (Faraday's law of induction)
- (ii) $\text{curl } \mathbf{H} = \frac{4\pi}{c} \mathbf{J} + \frac{1}{c} \frac{\partial \mathbf{D}}{\partial t}$ (Ampère's law)
- (iii) $\text{div } \mathbf{D} = 4\pi\rho$ (continuity)
- (iv) $\text{div } \mathbf{B} = 0$ (nonexistence of true magnetism)

Here c is the speed of light. We shall put these equations into the language of exterior forms. To this end, we set

$$\begin{aligned} \alpha &= (E_1 dx^1 + E_2 dx^2 + E_3 dx^3)(c dt) \\ &\quad + (B_1 dx^2 dx^3 + B_2 dx^3 dx^1 + B_3 dx^1 dx^2), \\ \beta &= -(H_1 dx^1 + H_2 dx^2 + H_3 dx^3)(c dt) \\ &\quad + (D_1 dx^2 dx^3 + D_2 dx^3 dx^1 + D_3 dx^1 dx^2), \\ \gamma &= (J_1 dx^2 dx^3 + J_2 dx^3 dx^1 + J_3 dx^1 dx^2) dt - \rho dx^1 dx^2 dx^3. \end{aligned}$$

from
(124) = $\{x^1, x^2, x^3, t\}$

$C = \text{constant}$
($C = 1 \text{ or } c$)

Equations (i) and (iv) become

$$d\alpha = 0.$$

Equations (ii) and (iii) become

$$d\beta + 4\pi\gamma = 0.$$

Applying d to this last equation yields

$$d\gamma = 0,$$

in vector notation

$$\text{div } \mathbf{J} + \frac{\partial \rho}{\partial t} = 0.$$

From the equation $d\alpha = 0$ one concludes, at least in any region of space-time which can be shrunk to a point, that there is a one-form λ such that

$$d\lambda = \alpha.$$

We introduce the vector potential \mathbf{A} and a scalar A_0 by writing

$$\lambda = A_1 dx^1 + A_2 dx^2 + A_3 dx^3 + A_0 c dt.$$

The equation $d\lambda = \alpha$ in vector form is

$$\begin{cases} \text{curl } \mathbf{A} = \mathbf{B} \\ \text{grad } A_0 - \frac{1}{c} \frac{\partial \mathbf{A}}{\partial t} = \mathbf{E}. \end{cases}$$

from Flanders, Differential Forms with Applications to the Physical Sciences

In free space, everything simplifies according to

$$\mathbf{E} = \mathbf{D}, \quad \mathbf{H} = \mathbf{B},$$

$$\mathbf{J} = 0, \quad \rho = 0$$

so that the Maxwell equations become

$$\begin{cases} \text{curl } \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{H}}{\partial t} & \text{div } \mathbf{E} = 0 \\ \text{curl } \mathbf{H} = \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t} & \text{div } \mathbf{H} = 0. \end{cases}$$

We introduce the Lorentz metric into 4-space whereby

$$dx^1, dx^2, dx^3, c dt$$

is an orthonormal basis:

$$(dx^i, dx^j) = \delta^{ij}, \quad (dx^i, c dt) = 0,$$

$$(c dt, c dt) = -1.$$

The signature is $3 - 1 = 2$.

According to the formulas of Section 2.7,

$$*(dx^1 dx^2) = -dx^3 (c dt), \quad \text{etc.},$$

$$*(dx^1 c dt) = dx^2 dx^3, \quad \text{etc.}$$

We see that

$$\begin{aligned} \alpha &= (E_1 dx^1 + \dots)(c dt) + (H_1 dx^2 dx^3 + \dots), \\ \beta &= -(H_1 dx^1 + \dots)(c dt) + (E_1 dx^2 dx^3 + \dots) \\ &= *\alpha. \end{aligned}$$

Consequently Maxwell's equations in free space are simply

$$\begin{cases} d\alpha = 0 \\ d*\alpha = 0. \end{cases}$$

We return to the general situation and refine our analysis by introducing one-forms:

$$\omega_1 = E_1 dx^1 + E_2 dx^2 + E_3 dx^3$$

$$\omega_2 = B_1 dx^2 dx^3 + B_2 dx^3 dx^1 + B_3 dx^1 dx^2$$

$$\omega_3 = H_1 dx^1 + H_2 dx^2 + H_3 dx^3$$

$$\omega_4 = D_1 dx^2 dx^3 + D_2 dx^3 dx^1 + D_3 dx^1 dx^2$$

$$\omega_5 = J_1 dx^2 dx^3 + J_2 dx^3 dx^1 + J_3 dx^1 dx^2.$$

These involve space variable differentials only. Now we interpret d' to denote the exterior derivative with respect to space variables only. We introduce $\partial/\partial t$ in this form

$$\frac{\partial}{\partial t}(\omega_1) = \dot{\omega}_1 = \dot{E}_1 dx^1 + \dots, \text{ etc.}$$

Now the Maxwell equations are

$$\begin{cases} d'\omega_1 = -\frac{1}{c}\dot{\omega}_2 \\ d'\omega_3 = \frac{4\pi}{c}\omega_5 + \frac{1}{c}\dot{\omega}_4 \\ d'\omega_2 = 0 \\ d'\omega_4 = 4\pi\rho dx^1 dx^2 dx^3. \end{cases}$$

The Poynting energy-flux vector \mathbf{S} is introduced by

$$\mathbf{S} = \left(\frac{c}{4\pi}\right) \mathbf{E} \times \mathbf{H}$$

that is

$$\left(\frac{c}{4\pi}\right) \omega_1 \wedge \omega_3 = S_1 dx^2 dx^3 + S_2 dx^3 dx^1 + S_3 dx^1 dx^2.$$

Poynting's theorem,

$$\left(\frac{1}{4\pi}\right) \dot{\mathbf{B}} \cdot \mathbf{H} + \mathbf{E} \cdot \mathbf{J} + \left(\frac{1}{4\pi}\right) \mathbf{E} \cdot \dot{\mathbf{D}} + \text{div } \mathbf{S} = 0,$$

follows from

$$\begin{aligned} d'(\omega_1 \wedge \omega_3) &= d'\omega_1 \wedge \omega_3 - \omega_1 \wedge d'\omega_3 \\ &= \left(-\frac{1}{c}\dot{\omega}_2\right) \wedge \omega_3 - \omega_1 \wedge \left(\frac{4\pi}{c}\omega_5 + \frac{1}{c}\dot{\omega}_4\right) \\ &= -\frac{1}{c}\dot{\omega}_2 \wedge \omega_3 - \frac{4\pi}{c}\omega_1 \wedge \omega_5 - \frac{1}{c}\omega_1 \wedge \dot{\omega}_4. \end{aligned}$$

For bodies at rest, one assumes $\mathbf{D} = \kappa\mathbf{E}$, $\mathbf{B} = \mu\mathbf{H}$ where the dielectric constant κ and the permeability μ are constant in time. Then Poynting's theorem becomes

$$-\frac{\partial u}{\partial t} = \text{div } \mathbf{S} + \mathbf{E} \cdot \mathbf{J}$$

where

$$u = \frac{1}{8\pi}(\kappa\mathbf{E}^2 + \mu\mathbf{H}^2)$$

is the *energy density* of the field. The quantity $\mathbf{E} \cdot \mathbf{J}$ is called the *therm chemical activity*.

4.7. Problems

1. Develop the formula for the Laplacian in cylindrical coordinates.
2. A complex matrix A is *unitary* if $AA^* = I$, where $A^* = \overline{A}^t$, the transpose conjugate of A . We call A skew-hermitian if $A^* + A = 0$. Discuss the connection between unitary and skew-hermitian matrices.
3. Show that e^A is orthogonal if A is skew-symmetric. Here

$$e^A = I + \sum_{n=1}^{\infty} \frac{A^n}{n!}$$

for the real matrix A .

4. Set up a frame and the structure equations for a sphere of radius r . Compute the curvatures.
5. Find Gaussian curvature of the surface of revolution obtained by revolving the curve

$$\begin{cases} x = \cos \theta + \ln \tan(\theta/2) \\ y = \sin \theta \\ \frac{\pi}{2} < \theta < \pi \end{cases}$$

about the x -axis.

6. Given a surface in the form $z = f(x, y)$, develop formulas for H and K in terms of f and its partial derivatives.
7. Let Σ be a surface. Let $\mathbf{e}_1, \mathbf{e}_2$ and $\mathbf{e}'_1, \mathbf{e}'_2$ be two moving frames of tangent vectors to Σ . Determine the relation between the corresponding ω and ω' and verify that $d\omega = d\omega'$.
8. Let $\mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3$ and $\mathbf{e}'_1, \mathbf{e}'_2, \mathbf{e}'_3$ be two moving frames in \mathbb{E}^3 . Set up the orthogonal matrix relating these frames and determine how the corresponding Ω and Ω' are related.